

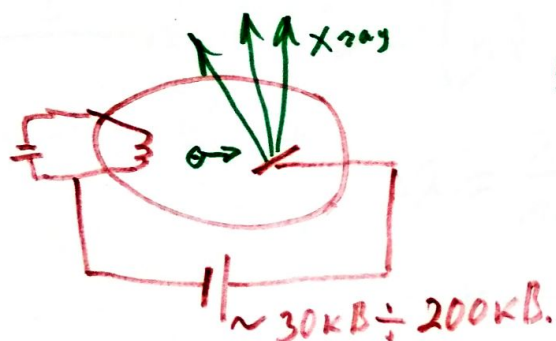
$$h\nu = A_0 + \frac{mv^2}{2}$$

$$1. N_3 \sim N_0 \sim I_H \sim \Phi$$

$$2. eU_3 = \frac{mv^2}{2} = h\nu - A_0$$

$$3. h\nu < A_0 - \text{нет!} \quad \nu_k = \frac{A_0}{h}$$

① В. Рентген 1895г X-ray $\lambda \approx 1 \text{ \AA} = 10^{-10} \text{ м}$

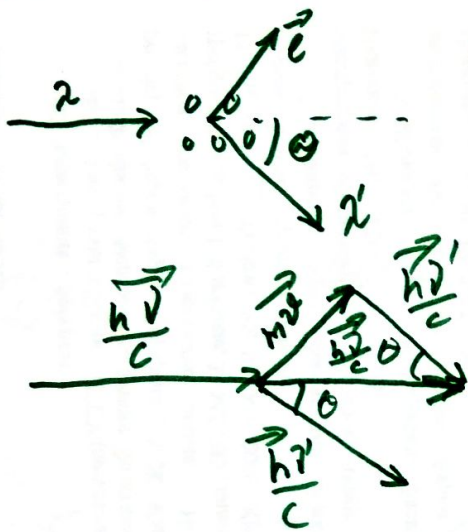


Дж. Дж. Томсон

$$\frac{mv^2}{2} = h\nu_{\max}$$



КОМПТОН 1923



$$\Delta\lambda = \lambda' - \lambda = 2K \sin^2 \frac{\theta}{2} \quad K = 0,0241 \cdot 10^{-10} \text{ m}$$

$$p_\gamma = \frac{h\nu}{c}$$

$$(m_e v)^2 = \left(\frac{h\nu}{c}\right)^2 + \left(\frac{h\nu'}{c}\right)^2 - 2\left(\frac{h\nu}{c}\right)\left(\frac{h\nu'}{c}\right)\cos\theta$$

$$h\nu + m_e c^2 = h\nu' + m_e c^2$$

$$m = \frac{m_0}{\sqrt{1 - \frac{v^2}{c^2}}}$$

$$\Delta\lambda = \frac{2h}{m_e c} \sin^2 \frac{\theta}{2}$$

$$K = \frac{h}{m_e c}$$

Вавилов, Харитон.



Давление света Н. Лебедев 1898 г.

$$P = \frac{I}{c} (1 + \rho)$$

Боте.



В Луи де Бройль 1924

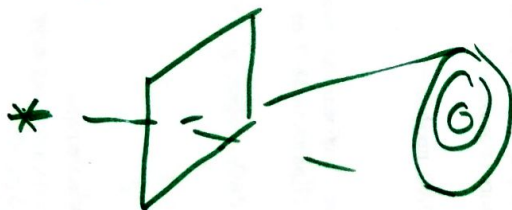
$$E = h\nu = mc^2 = pc \Rightarrow \lambda_B = \frac{c}{\nu} = \frac{h}{p} = \frac{2\pi\hbar}{p} -$$

$$\hbar = \frac{h}{2\pi}$$

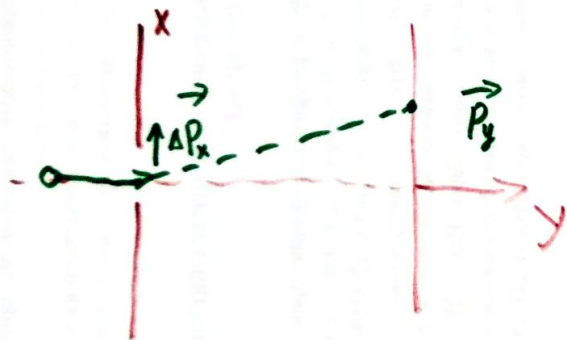
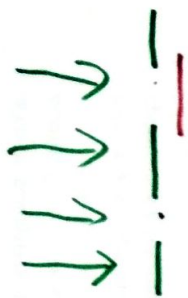
1927 Делюсон, Дхермер.

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1928 Др. П. Томсон. Тартаковский



1948 Рабрикант 10^{14}



Гейзенберг 1927

неопределённостей

$$\begin{cases} \Delta x \Delta p_x \geq \frac{\hbar}{2} \\ \Delta y \Delta p_y \geq \frac{\hbar}{2} \\ \Delta z \Delta p_z \geq \frac{\hbar}{2} \end{cases}$$

$\hbar, \frac{h}{2}, h$

$$[x, p] = \frac{h}{2\pi i}$$

$$[E, t] = \hbar$$

$$\Delta E \Delta t \geq \frac{\hbar}{2}$$

канонически